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2.

We consider a Hamiltonian of the form:

$$\hat{H} = \hat{H}^{(0)} + \lambda \hat{H}^{(1)}$$

Where:

- $\hat{H}^{(0)}$ is the unperturbed Hamiltonian.
- $\hat{H}^{(1)}$ is the perturbation.
- λ is a small parameter.

The time-independent Schrödinger equation becomes:

$$\hat{H}|\psi_n\rangle = E_n|\psi_n\rangle$$

Suppose the unperturbed energy level $E^{(0)}$ is **d-fold degenerate**, meaning there are d orthonormal eigenstates $\{|\psi_a^{(0)}\rangle\}_{a=1}^d$ such that:

$$\hat{H}^{(0)}|\psi_a^{(0)}\rangle = E^{(0)}|\psi_a^{(0)}\rangle$$

To find the first-order corrections, we diagonalize the perturbation $\hat{H}^{(1)}$ **within the degenerate subspace**:

$$H_{ab}^{(1)} = \langle \psi_a^{(0)} | \hat{H}^{(1)} | \psi_b^{(0)} \rangle$$

We solve the eigenvalue problem:

$$\sum_{b=1}^d H_{ab}^{(1)} c_b = E^{(1)} c_a$$

Cb-mis see on?

Kuidas see võrrand saadi?

10.

Derive and explain!

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Photon Emission Rate (Transition $n \rightarrow m$)

The Fermi's Golden Rule gives the transition rate from an initial state $|n\rangle$ to a final state $|m\rangle$ due to a perturbation $\hat{H}^{(1)}(t)$:

$$\Gamma_{n \rightarrow m} = \frac{2\pi}{\hbar} |\langle m | \hat{H}^{(1)} | n \rangle|^2 \rho(E_m)$$

Where:

- $\Gamma_{n \rightarrow m}$ is the **transition rate** (photons per second).
- $\langle m | \hat{H}^{(1)} | n \rangle$ is the **matrix element** of the perturbation between the states.
- $\rho(E_m)$ is the **density of final states** at energy E_m .

Physical Interpretation

- This formula tells us how likely it is per unit time for a system in state $|n\rangle$ to transition to $|m\rangle$, emitting a photon of energy $\hbar\omega = E_n - E_m$.
- The **number of photons emitted per second** is directly proportional to the square of the transition matrix element and the density of available final states.

15.

1. Schrödinger Equation

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- **Type:** Non-relativistic quantum mechanics
- **Form:**

$$i\hbar \frac{\partial}{\partial t} \psi(\mathbf{r}, t) = \hat{H} \psi(\mathbf{r}, t)$$

- **Describes:** Particles moving at speeds much less than the speed of light.
- **Wavefunction:** Scalar function ψ
- **Does not include:** Spin or relativistic effects.
- **Limitation:** Cannot describe particles moving near the speed of light or account for antimatter.

2. Dirac Equation

- **Type:** Relativistic quantum mechanics
- **Form:**

How about an explanation of the parameters?

$$(i\hbar \gamma^\mu \partial_\mu - mc) \psi = 0$$

- **Describes:** Particles like electrons at relativistic speeds.
- **Wavefunction:** Four-component spinor (accounts for spin-1/2 particles).
- **Includes:**
 - Special relativity
 - Spin naturally
 - Predicts antimatter (e.g., positrons)
- **Reduces to Schrödinger equation** in the low-energy limit.